## Statistical Orbit Determination



Lecture 6 - Coordinate Systems and Time
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## Recap

- Lecture 5 - Notes posted here
- Perturbed Motion
- Questions
- Post them to YouTube page


## Agenda

- Coordinate Systems and Time: Introduction
- Precession and Nutation
- Earth Rotation and Time
- Earth-Fixed and Topocentric Systems
- Transformation between ECF and ECI
- Orbit Accuracy
- Assigned Problems


## Reference Frame - Rotation and Time (1/5)

- Rotation, $\omega_{\oplus}$
- Simplified model is fixed in direction and magnitude
- Conical CCW motion, full revolution in 430 days, Chandler period
- The polar motion is confined to 0.6 arcsec over decadal time scale
- Changes in magnitude and length of day
- Change in $\omega_{\oplus}$ means change period required for rev of Earth $w / r$ stars
- Not w/r to the sun
- Period change expressed in form of Universal Time, $\Delta$ (UT 1 )
- Difficult to predict period change
- Regular observation and reported by International Earth Rotation Service


## Reference Frame - Rotation and Time (2/5)

- Estimation of $\Delta$ (UT1)
- Separation of $\Delta$ (UTI) from orbit node, $\Omega$, complicates estimation
- Polar motion and satellite state can be used (GPS and Lageos)
- Errors in $\Omega$ absorbed in estimation of $\Delta$ (UT1)
- Cannot provide reliable long-term estimates of $\Delta$ (UT1)
- Time is different
- Event uniquely identified by local time or Universal Time
- Epoch J2000.0 is January 1, 2000, 12 hours
- Julian Date (JD)
- Measured from 4713 b.c. and day begins at noon
- J2000.0 is JD 2451545.0 days
- January 1, 2000, 0 hours (midnight) would be JD 2451544.5


## Reference Frame - Rotation and Time (3/5)

- Modified Julian Date (MJD)
- JD minus 2400000.5
- J2000.0 is MJD 51544.5 day
- Reference time necessary because independent variable
- Every observation is time tagged using a reference clock
- Reference clocks use oscillators
- Typically crystal oscillators used
- Affected by temperature and some temporal characteristics
- High accuracy uses atomic frequencies



## Reference Frame - Rotation and Time (4/5)

- Various time systems
- Terrestial Dynamical Time (TDT)
- Used for Earth satellites
- Also known as terrestrial time (TT)
- Barycentric Dynamical Time (TDB)
- Used for solar system applications
- International Atomic Time (TAI)
- Both TDT and TDB related to TAI at a specified epoch
- Based on cesium atomic clocks
- Universal Time
- Measure of time that is basis for all civil time-keeping
- Coordinated Universal Time (UTC) derived from TAI where UT1 = UTC $+\Delta$ (UT1)
- UTC maintained by NIST and USNO
- UTC(USNO) and TAI based on ensemble of cesium oscillators and hydrogen masers


## Reference Frame - Rotation and Time (5/5)

- Difference between time systems
- UTC needs leap seconds to maintain synchronization
- UTC and $\Delta$ (UT1) require leap seconds for synchronization within $\pm 0.9$ seconds
- Constant difference between TAI and TT
- TT - TAI $=32.184$ seconds
- TDB and TT difference is periodic function, relativistic effect
- GPS time (GPS-T) related to TAI
- Leap second adjustments applied to UTC are not used in GPS-T
- Current leap second separation is 37.0 seconds as of Jan 1, 2017


## Reference Frame - EF and Topocentric (1/3)

- Earth-fixed and topocentric systems
- Earth-fixed frame not defined since not a rigid body
- Mass deformation of luni-solar gravity changes coordinates on surface
- True Earth-fixed frame does not exist
- Terrestrial Reference Frame (TRF)
- Origin is coincident with center of mass
- "Attached," $x$-axis approx. coincident with Greenwich meridian
- z-axis approx. coincident with $\omega_{\oplus}$
- International Earth Rotation Service TRF or ITRF
- WGS-84, used with many GPS applications



## Reference Frame - EF and Topocentric (2/3)

- Ellipsoid of revolution not sphere
- Spherical coordinates used to describe gravitational potential

$$
\begin{gathered}
x=r \cos \phi \cos \lambda \\
y=r \cos \phi \sin \lambda \\
z=r \sin \phi
\end{gathered}
$$

- $\phi$, geocentric latitude, $\lambda$, longitude, and $r$ is magnitude of pos vector
- Alternate set used with ellipsoid
- Geodetic latitude, $\phi^{\prime}$, longitude, $\lambda$, and height above ellipsoid, $h$

$$
\begin{gathered}
x=\left(N_{h}+h\right) \cos \phi^{\prime} \cos \lambda \\
y=\left(N_{h}+h\right) \cos \phi^{\prime} \sin \lambda \\
z=\left(N_{h}+h-\bar{e}^{2} N_{h}\right) \sin \phi^{\prime}
\end{gathered}
$$

- Where eccentricity of the elliptical cross-section is

$$
\begin{gathered}
\bar{e}^{2}=\bar{f}(2-\bar{f}) \\
N_{h}=\frac{R_{e}}{\left(1-\bar{e}^{2} \sin ^{2} \phi^{\prime}\right)^{1 / 2}} \\
\bar{f}=\frac{R_{e}-R_{p}}{R_{e}} \\
x^{2}+y^{2}+\left(\frac{R_{e}}{R_{p}}\right)^{2} z^{2}=R_{e}^{2}
\end{gathered}
$$

## Reference Frame - EF and Topocentric (3/3)

- Topocentric
- Northward, Eastward, local vertical point on a surface
- $x_{t}$ eastward, $y_{t}$ northward, $z_{t}$ local vertical
- Earth-fixed in terms of topocentric system is

$$
\bar{r}_{t}=T_{t}\left(\bar{r}-\bar{r}_{s}\right)=T_{t} \bar{\rho}
$$

$$
T_{t}=\left[\begin{array}{ccc}
-\sin \lambda & \cos \lambda & 0 \\
-\sin \phi \cos \lambda & -\sin \phi \sin \lambda & \cos \phi \\
\cos \phi \cos \lambda & \cos \phi \sin \lambda & \sin \phi
\end{array}\right]
$$

- Elevation and Azimuth

$$
\begin{aligned}
& \sin (E l)=\frac{z_{t}}{r_{t}} \quad-90^{\circ} \leq E l \leq 90^{\circ} \\
& \sin (A z)=\frac{x_{t}}{r_{x y}} \quad 0 \leq A z \leq 360^{\circ} \\
& \cos (A z)=\frac{y_{t}}{r_{x y}}
\end{aligned}
$$

## Reference Frame - Transform ECF and ECI

- Transform between ECF and ECl systems
- Transformation matrix between ECF to J2000 system is complex
- Must consider precession, nutation, polar motion, and UT 1

$$
T_{X Y Z}^{x y Z}=W S^{\prime} N P
$$

- Where the transformation is from J2000 to ECF,
- $P$ applies to precession from epoch to current time
- $N$ applies nutation at current time
- $S^{\prime}$ applies rotation to account for true sidereal time
- $W$ applies polar motion to align the $z$ axis (true pole) with the pole of the ECF system


## Orbit Accuracy

## - Orbit accuracy

- Accuracy of solution of EOM with parameters in model of forces
- General perturbation technique concerned with small parameters
- Special perturbation technique concerned with step size
- Solution technique accuracy
- Error introduced in the solution of the equations of motion by the solution technique
- No consideration given to the accuracy of the parameters in the equations
- Force model accuracy
- Parameters in force models and modeling of forces most significant error source
- All parameters in force model have been determined by some means
- Specification of requirement
- Equations of motion distinctly different if orbit must be determined with cm vs km accuracy


Practice problems: The Orbit Problem PREDICT THE ORBIT

## Assigned Problems - Overview

- You are given three problems involving orbital motion. They have been picked to ensure you have a sufficient understanding of orbital mechanics before proceeding. The problems resemble numbers $4,5,6,10,11$, and 12 from the textbook.
- These problems should be complete by Friday, February 8.


## Assigned Problems - Problem 1

- Given the following position and velocity of a satellite
- Expressed in a non-rotating geocentric coordinate system

|  | Position (m) | Velocity $(\mathrm{m} / \mathrm{s})$ |
| ---: | ---: | ---: |
| X | 7088580.789 | -10.20544809 |
| Y | -64.326 | -522.85385193 |
| Z | 920.514 | 7482.075141 |

a) Determine the six orbital elements ( $\mathrm{a}, \mathrm{e}, \mathrm{i}, \Omega, \omega, \mathrm{M}_{0}$ )
b) Assuming $X_{0}$ is given and two-body motion, predict position and velocity at $t=$ $3,000 \mathrm{sec}$. Determine flight path angle at this time.
c) Determine the latitude and longitude of the subsatellite point for $t=3,000 \mathrm{sec}$ if $\alpha_{G}$ at $t=0$ is 0 . Assume the $Z$ axis of the nonrotating system is coincident with the $Z$ axis of the rotating system.

## Assigned Problems - Problem 2 (1/2)

## - Orbit of CRISTA-SPAS-2

- Cryogenic Infrared Spectrometers and Telescopes for the Atmosphere

The joint venture of DLR and NASA, the small free-flying satellite contains three telescopes, four spectrometers, and a GPS receiver on-board. It is deployed from the shuttle Discovery on STS-85 in August 1997. Using on-board navigation, the receiver measurements are processed in an Earth-centered, Earth-fixed coordinate system.

| August 18, 1997 |  |  |
| :---: | :---: | ---: |
| GPS-T (hrs:min:sec) | $00: 00: 0.000000$ | $00: 00: 03.000000$ |
| $x$ | 3325396.441 | 3309747.175 |
| $y$ | 5472597.483 | 5485240.159 |
| $z$ | -2057129.050 | -2048664.333 |
|  | August 19,1997 |  |
| GPS-T (hrs:min:sec) | $00: 00: 0.000000$ | $00: 00: 03.000000$ |
| $x$ | 4389882.255 | 4402505.030 |
| $y$ | -4444406.953 | -4428002.728 |
| $z$ | -2508462.520 | -2515303.456 |

## Assigned Problems - Problem 2 (2/2)

a) Demonstrate that the node location is not fixed in space and determine an approximate rate of node change (degrees/day) from these positions.
Compare the node rate with the value predicted by

$$
\dot{\Omega}=-\frac{3}{2} J_{2} \frac{n}{\left(1-e^{2}\right)^{2}}\left(\frac{a_{e}}{a}\right)^{2} \cos i
$$

b) Determine the inclination of CRISTA-SPAS-2 during the first 3-sec interval and the last $3-\mathrm{sec}$ interval.

Comment: The position vectors determined by GPS in this case are influenced at the 100-meter level by Selective Ability, but the error does not significantly affect this problem.

## Assigned Problems - Problem 3 (1/2)

## - GLONASS

- Russia's answer for American GPS

Given a set of initial conditions for a high-altitude GLONASS satellite, numerically integrate the equations of motion for one day.
a) Assuming the satellite is influenced by $J_{2}$ only, derive the equations of motion in non-rotation coordinates. Assume the nonrotating $Z$ axis coincides with the Earth-fixed $z$ axis.
b) During the integration, compute the Jacobi constant and the $Z$ component of the angular momentum. Are these quantities constant?
c) Plot the six orbital elements as a function of time.
d) Identify features similar to and different from Fig. 2.3.5


## Assigned Problems - Problem 3 (2/2)

e) Compare the node rate predicted by

$$
\dot{\Omega}=-\frac{3}{2} J_{2} \frac{n}{\left(1-e^{2}\right)^{2}}\left(\frac{a_{e}}{a}\right)^{2} \cos i
$$

with a value estimated from (c).
f) Compare the amplitude of the semimajor axis periodic term with

$$
a(t)=\bar{a}+3 \bar{n} \bar{a} j_{2}\left(\frac{a_{e}}{\bar{a}}\right)^{2} \sin ^{2} \frac{\bar{l}(\cos (2 \omega+2 M))}{2 \dot{\omega}_{s}+2 \dot{M}_{s}}
$$

g) Plot the ground track. Does the ground track repeat after one day?

| a | e | i | $\Omega$ | $\omega$ | $M_{0}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 25500.0 km | 0.0015 | 63 deg | -60 deg | 0 deg | 0 deg |

